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Armando: The Final Subcritical Experiment in the Series

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Introduction

Armando was the final subcritical experiment (SCE) in the Stallion series. A series of experiments (Vito/Etna, Mario, Rocco, and Armando) were designed to study the high-explosive (HE) driven properties of cast and wrought plutonium; the materials being representative of the materials produced via the different processes employed at Rocky Flats and LANL. Specific properties investigated include production, spall features, and surface temperatures. Ejecta is defined as particulate matter "ejected" from the surface of a solid when a strong shock is applied to the surface. Spall is a general term for bulk material failure near the surface created by a strong shock interacting with the surface. Both of these phenomena depend upon the material properties: material strength, grain size, impurities, etc. The strength and temporal profile of the shock pressure. The surface temperature is an important constraint upon the final state of the shock-driven metal that is required for a full understanding of the behavior of the material.

Vito/Etna was a joint experiment conducted with the Aldermaston Weapons Research Establishment (AWE) located in the United Kingdom that concentrated on ejecta. Rocco and Mario were separate cast and wrought experiments that examined the phenomena of ejecta via a suite of point diagnostics. These diagnostics provided either specific time or position data at a single spatial point or a continuous time record of some material property at a spatial point. Armando was designed to complement and extend the measurements of Rocco and Mario by combining the two experiments into a single package. The radiographing the behavior at two separate times along equivalent lines. The paraphrased paradigm is: a picture is worth a thousand pins.

Diagnostics and Package Composition

HE diagnostics equivalent to those used on Rocco and Mario were implemented to ensure identical HE performance. These consisted of a series of shorting switches, a microwave interferometer strip laid out symmetrically on the HE package to measure detonation times and velocities. Point VISARs (Velocity Interferometer System) and optical pyrometers (that provide a measure of the surface temperature) were also implemented to verify equivalent behavior of the surface properties. A diagnostic for Armando was x-ray radiography along two equivalent axes. The Physics packages identical to the Rocco and Mario packages were combined with the HE package (6-sides or High-Energy X-ray) vertically separated with the free surface of one on another. This geometry allows for exactly equivalent radiographs to be taken of two materials at the same time in their evolution. The third axis of the HE package is for VISAR/pyrometry access.

within a 3 ft diam (inside diameter) containment vessel. This vessel and that houses the scintillator and camera system are placed within a "zero" a large bulkhead completely sealing off the end of the U1a.05 drift. Thin windows in the bulkhead and vessel allow the x-rays to pass through the containment vessel with minimal attenuation. Originally designed so as to field the experiment within a containment vessel allows for the room for multiple experiments and provides multiple redundant containm

The Cygnus sources extend down the drift externally to the zero room. T of a Marx bank system contained in large oil-filled tanks that pulse-charge forming lines (PFL). The output of the PFL is a short pulse (~ 60 ns), large MV) electrical pulse that propagates down an 8 in. diam, water-filled, co line. This electrical pulse is coupled into the inductive voltage adder (IVA voltage in parallel to produce a 2.25 MV, low-impedance drive pulse for t diode. This last stage of voltage addition is accomplished in a high-vacuum diode operation. The expertise of Sandia National Laboratories (SNL) and Sciences Division (PSD) were instrumental in applying the technology de realize a robust and flexible pulse-powered driver capable of operating re underground.

Figure 2. Layout of Cygnus x-ray sources in U1a.05 drift. Containment vessel and camera box are not shown in this view.

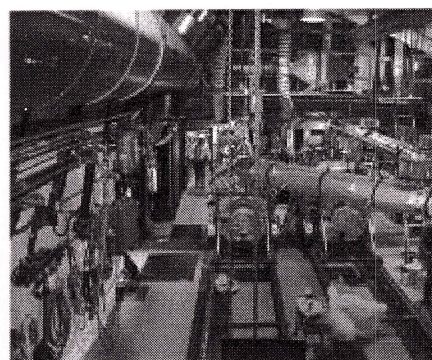
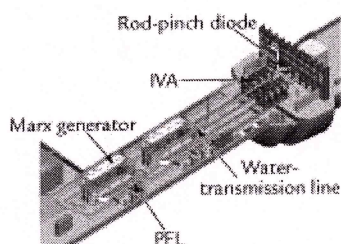


Figure 3. A view of the U1a.05 drift. The "zero" bulkhead is seen with the two IVA structures fed water-filled coaxial transmission lines in the fore

Radiography and Detector Systems

The radiography employed on Armando represents a significant advance in performance of medium-energy radiography. It has been the result of a Laboratory effort involving LANL, SNL, Bechtel Nevada (BN), the Naval Research Laboratory (NRL), AWE, Titan/PSD, and Mission Research Corporation (MRC). Many factors have been combined to lead to this advance in performance, but perhaps the most significant has been the effective realization of the rod-pinch diode originally developed by NRL. The rod-pinch has a similar geometry to standard x-ray diodes in use in industrial radiography sources for several decades. However, researchers at NRL realized that by using a low impedance (Z), the diode would transition from classic space charge limited (SCL) flow into magnetically limited (ML) flow whereby the electrons would be transplanted from the central anode rod and then "pinch" producing a very bright, small

reproducible manner.

The detector system is equally innovative. It combines technologies developed for Dual-Axis Radiographic Hydrodynamic Test (DARHT) and proton radiography to create a very high-resolution imaging system. It functions by converting the x-rays transmitted through the experimental package into visible light in a tiled LSO (lutetium oxyorthosilicate) scintillator. The light produced is transported by a low-f/# lens system to a LN₂-cooled charged-coupled-device (CCD) chip that captures and records the image. In order to preserve maximum resolution, the combined CCD camera system is not gated; all time resolution derives from the flash nature of the illuminating x-ray pulse. While providing high resolution, this technique introduces a risk to the experiment, in that the camera combination must be maintained in a light tight configuration through the detonation and long enough thereafter (~ 90 s) for the information to be transferred from the CCD camera system to a remote data logging computer.

Results

The results of Armando have provided valuable data for stockpile stewardship diagnostics demonstrated identical performance of the HE detonation in Rocco, Mario, and Armando cast and wrought. The VISAR measurement of the measured surface velocity was reproduced within error bars. The radiographic data provided detailed subsurface data on the spalled material with a resolution and precision previously unobtainable. Radiographic data was obtained at two times allowing comparison with VISAR data. The inferred velocity was in excellent agreement, further enhancing the accuracy of the results.

Acknowledgment

A great many people in many divisions (Physics, Dynamic Experimentation, Sciences and Applications, Material Science and Technology, Applied Physics, Materials Technology, Health, Safety, and Radiation Protection, and Earth and Environmental Sciences) at LANL worked in a very productive partnership with PSD, NRL, MRC, and BN to develop the technology and execute the Armando project. It is not possible to call out every individual in this format, but their efforts and dedication to the project are deeply appreciated.

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